















Environment

# Statistical Climate Reconstruction Modelling in the EUSTACE Project

Finn Lindgren (finn.lindgren@ed.ac.uk)

#### The University of Edinburgh, Scotland

with Colin Morice. John Kennedy, and the EUSTACE team. David Bolin, Haavard Rue, Daniel Simpson, Elias Krainski

Modern Statistical and Machine Learning Approaches for High-Dimensional Compound Spatial Extremes

BIRS-IMAG, Granada, 7–12 May 2023

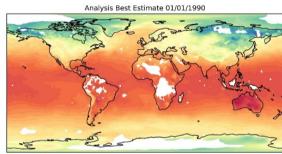


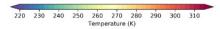


# **EUSTACE ANALYSIS**

Combines in-situ and satellite data sources to derive daily air temperatures across the globe with quantified uncertainties.

- Daily mean air temperature (2 m) estimates from the midlate 19th century at ¼ degree resolution.
- Observational dataset for use in climate monitoring, services and research.
  - Quantify bias and uncertainty arising from observational sampling (in space and time);
  - Quantify uncertainty from instrumental effects/network changes.
- Higher resolution daily gridded analyses for regional climate
  - Combine in situ and remote sensing data to support high resolution analysis.
  - Absolute temperature rather than anomaly product.





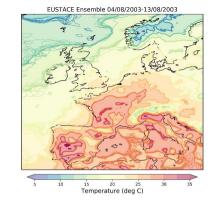


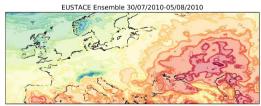


#### **ENSEMBLE ANALYSIS**

- Samples drawn from joint posterior distribution of temperature and bias variables.
- Temperature model samples projected onto analysis grid.
- Spatial/temporal correlation in analysis errors is encoded into the ensemble.
- Summary statistics can be derived from the ensemble.
   Expected value, total uncertainty and observation constraint information also available.





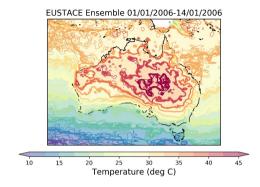


Temperature (deg C)

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Temperature (deg C)

# MULTI-SCALE ANALYSIS MODEL

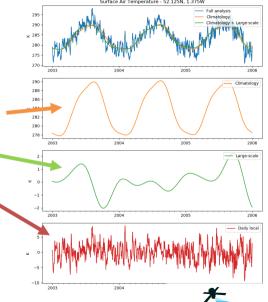
Statistical model for temperature variations and different scales (space and time):

- Climatological variation: local seasonal cycle with effects of latitude, altitude and coastal influence.
- Large-scale variation: Slowly varying climatological mean temperature field.
- Daily Local: daily variability associated with weather.

Simultaneously estimates observational biases of known bias structures:

• e.g. satellite biases, station homogenisation.

#### **Central England Temperature Decomposition**

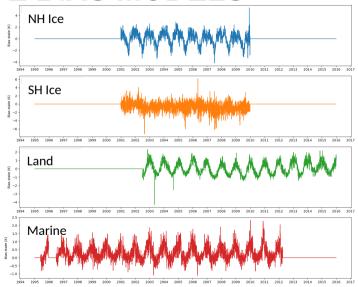


EUSTAC



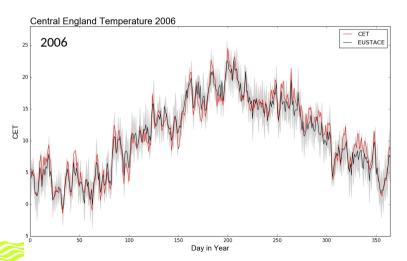
# SATELLITE BIAS MODELS

- Simplified model of known error structures in satellite air temperature retrievals:
  - Global/hemispheric systematic bias covariates.
  - Daily estimates of spatially varying bias as a spatial random field.
- Estimated jointly with daily temperature variability.

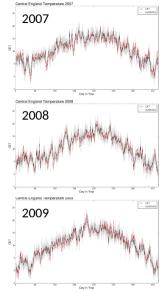




#### COMPARING EUSTACE WITH CENTRAL ENGLAND TEMPERATURE



Met Office Hadley Centre





# MULTI-SCALE ANALYSIS MODEL

Statistical model for temperature variations and different scales (space and time):

- Climatological variation: local seasonal cycle with effects of latitude, altitude and coastal influence.
- Large-scale variation: Slowly varying climatological mean temperature field. Station homogenisation.
- Daily Local: daily variability associated with weather.
   Satellite retrieval biases

Simultaneously estimates observational biases of known bias structures:

• e.g. satellite biases, station homogenisation.

Processed on STFC's LOTUS cluster www.jasmin.ac.uk:

- Largest solves processed on 20 core/256GB RAM node.
- Highly parallel observation pre-processing.



Element	Resolution	N Variables
Seasonal	Bimonthly x 1° SPDE	245,772
Slow-scale*	5 year x 5° SPDE	107,604
Latitude	0.5° latitude SPDE	721
Altitude	(0.25° grid)	1
Coastal	(0.25° grid)	1
Grand mean	Analysis mean	1

Element	Resolution	N Variables
Large-scale	3 monthly x 5° SPDE	1,752,408
Station bias	NA	82,072

Element	Resolution	N Variables per day
Daily local	~0.5 degree SPDE	162,842
Satellite bias (marine)	Global	1
Satellite bias (land)	Global + 2.5 degree SPDE	1+40,962
Satellite bias (ice)	Hemispheric + 2.5 degree SPDE*	2 + 40,962



# **GAMs** and general kriging

Linear GAMs with GPs on space and covariates:

$$\eta_i = \sum_k f_k(z_{ik}) + u(\mathbf{s}_i),$$

each  $f_k(\cdot)$  and  $u(\cdot)$  represented with basis expansions with jointly Gaussian coefficients x.

- lacktriangle Linear observations with additive Gaussian observation noise:  $y=\eta+\epsilon=Ax+\epsilon$
- Covariance kriging

$$egin{aligned} oldsymbol{\Sigma}_{oldsymbol{y}} &= oldsymbol{A} oldsymbol{\Sigma}_{oldsymbol{x}} oldsymbol{A}^ op oldsymbol{\Sigma}_{oldsymbol{x}} oldsymbol{A}^ op oldsymbol{\Sigma}_{oldsymbol{x}}^{-1} (oldsymbol{y} - oldsymbol{A} oldsymbol{\mu}) \ &= oldsymbol{\mu} + oldsymbol{\Sigma}_{oldsymbol{x}} oldsymbol{A}^ op oldsymbol{\Sigma}_{oldsymbol{x}}^{-1} (oldsymbol{y} - oldsymbol{A} oldsymbol{\mu}) \end{aligned}$$

Precision kriging

$$\begin{aligned} Q_{x|y} &= Q_x + A^\top Q_\epsilon A \\ \mathsf{E}(x|y) &= \mu + Q_{x|y}^{-1} A^\top Q_\epsilon (y - A\mu) \end{aligned}$$





# Observation level covariance vs latent level precision

Covariance kriging: linear solve with a  $\Sigma$ ,  $\Sigma_{ij} = \text{Cov}(y_i, y_j)$  Vecchia approximation:

 $\mathbf{\Sigma}^{-1} pprox \mathbf{L} \mathbf{L}^{\top}$  for a given observation ordering, and sparse lower triangular L with given sparsity pattern;  $p(\mathbf{y}|\mathbf{\theta}) pprox p(y_1) \prod_{i=2}^n p(y_i|\mathbf{y}_{G_i}), G_i \subseteq \{1,\ldots,i-1\},$ 

$$\boldsymbol{a}^{\top} \boldsymbol{\Sigma}^{-1} \boldsymbol{a} \approx \sum_{i} (\sum_{j \in G_i} a_i L_{ij})^2$$

L obtained sequentially from  $\Sigma$  for each observation.

Precision kriging: linear solve with a  $m{Q}, Q_{ij} = \operatorname{Prec}(x_i, x_j | m{y})$ 

 $m{Q} = m{L} m{L}^ op$  for a given latent variable ordering, and sparse lower triangular L with the sparsity from  $m{Q}$  plus Cholesky infill.

The prior  $Q_x$  for SPDE process components is obtained via a local Finite Element construction, giving the model in a chosen finite function space closest to the full model.





## **Example model: Matérn driven heat equation on the sphere**

The iterated heat equation is a simple non-separable space-time SPDE family:

$$\left[\phi \frac{\partial}{\partial t} + (\kappa^2 - \Delta)^{\alpha_s/2}\right]^{\alpha_t} x(\mathbf{s}, t) dt = d\mathcal{E}_{(\kappa^2 - \Delta)^{\alpha_e}}(\mathbf{s}, t)/\tau$$

For constant parameters,  $x(\mathbf{s},t)$  has spatial Matérn covariance (for each t) in a Matérn-Whittle sense on  $\mathbb{S}^2$ .

#### Discrete domain Gaussian Markov random fields (GMRFs)

 ${m x}=(x_1,\ldots,x_n)\sim \mathcal{N}({m \mu},{m Q}^{-1})$  is Markov with respect to a neighbourhood structure  $\{\mathcal{N}_i,i=1,\ldots,n\}$  if  $Q_{ij}=0$  whenever  $j\neq \mathcal{N}_i\cup i$ .

Project the SPDE solution space onto local basis functions: random Markov dependent basis weights (Lindgren et al. 2011).

A finite element approximation has structure

$$x(s,t) = \sum_{i,j} \psi_i^{[s]}(s) \psi_j^{[t]}(t) x_{ij}, \quad x \sim \mathcal{N}(\mathbf{0},Q^{-1}), \quad Q = \sum_{k=0}^{\alpha_t + \alpha_s + \alpha_e} \boldsymbol{M}_k^{[t]} \otimes \boldsymbol{M}_k^{[\mathbf{s}]}$$
 even, e.g., if the spatial scale parameter  $\kappa$  is spatially varying.

# Classic and compact INLA methods ( $\sim$ description)

Laplace approximation at the conditional posterior mode  $x^*$ , and uncertainty integration:

$$p(\boldsymbol{\theta}|\boldsymbol{y}) \propto \frac{p(\boldsymbol{\theta})p(\boldsymbol{x}|\boldsymbol{\theta})p(\boldsymbol{y}|\boldsymbol{\theta},\boldsymbol{x})}{p(\boldsymbol{x}|\boldsymbol{\theta},\boldsymbol{y})}\bigg|_{\boldsymbol{x}=\boldsymbol{x}^*} \approx \frac{p(\boldsymbol{\theta})p(\boldsymbol{x}|\boldsymbol{\theta})p(\boldsymbol{y}|\boldsymbol{\theta},\boldsymbol{x})}{p_G(\boldsymbol{x}|\boldsymbol{\theta},\boldsymbol{y})}\bigg|_{\boldsymbol{x}=\boldsymbol{x}^*} = \widehat{p}(\boldsymbol{\theta}|\boldsymbol{y})$$
$$p(x_i|\boldsymbol{y}) = \int p(x_i|\boldsymbol{\theta},\boldsymbol{y})p(\boldsymbol{\theta}|\boldsymbol{y}) d\boldsymbol{\theta} \approx \sum_k \widehat{p}(x_i|\boldsymbol{\theta}^{(k)},\boldsymbol{y})\widehat{p}(\boldsymbol{\theta}^{(k)}|\boldsymbol{y})w_k = \widehat{p}(x_i|\boldsymbol{y})$$

- lacksquare Let  $\widehat{m{\mu}} = \mathsf{E}(m{x}|m{ heta},m{y})$  and  $m{Q}_{\epsilon} = abla_x 
  abla_x^ op \log p(m{y}|m{ heta},m{x}^*)$
- $\begin{array}{c} \blacktriangleright \quad \text{Classic method: Laplace approximation of each } \widehat{p}(x_i|\pmb{\theta},\pmb{y}), \text{ and} \\ \left\{ \begin{bmatrix} \pmb{A}\pmb{x} \\ \pmb{x} \end{bmatrix} | \pmb{\theta}, \pmb{y} \right\} \sim \mathcal{N} \left( \begin{bmatrix} \pmb{A}\widehat{\pmb{\mu}} \\ \widehat{\pmb{\mu}} \end{bmatrix}, \begin{bmatrix} \pmb{Q}_\epsilon + \delta \pmb{I} & -\delta \pmb{A} \\ -\delta \pmb{A}^\top & \pmb{Q}_x + \delta \pmb{A}^\top \pmb{A} \end{bmatrix}^{-1} \right), \text{ with } \delta \gg 0 \end{array}$
- $\qquad \qquad \textbf{Compact method: Variational approximation of } \widehat{p}(\boldsymbol{x}|\boldsymbol{\theta},\boldsymbol{y}), \text{ and } \\ \{\boldsymbol{x}|\boldsymbol{\theta},\boldsymbol{y}\} \sim \mathcal{N}\left(\widehat{\boldsymbol{\mu}},[\boldsymbol{Q}_x + \boldsymbol{A}^{\top}\boldsymbol{Q}_{\epsilon}\boldsymbol{A}]^{-1}\right)$





# Before satellites you had to go measure in person







## Hydrology lab from the 1925-27 Antarctic ocean expedition





"The Discovery", Dundee, Scotland (Photos: Finn Lindgren, August 2022)



#### What's that in the corner?

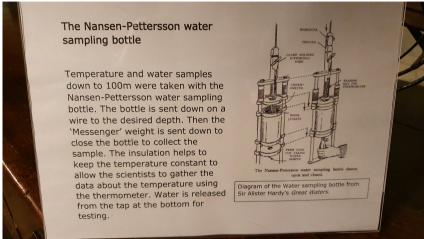




"The Discovery", Dundee, Scotland (Photos: Finn Lindgren, August 2022)



## It's a Nansen-Pettersson water sampling bottle!







# Station observation & homogenisation model

#### Daily mean air temperature measurements

For station k at day  $t_i$ ,

$$y_m^{k,i} = T_m(\mathbf{s}_k, t_i) + \sum_{i=1}^{J_k} H_j^k(t_i) e_m^{k,j} + \epsilon_m^{k,i},$$

where  $H^k_j(t)$  are temporal step functions,  $e^{k,j}_m$  are latent bias variables, and  $\epsilon^{k,i}_m$  are independent measurement and discretisation errors.

#### Daily mean/max/min

For station 
$$k$$
 at day  $t_i, y_m^{k,i} = T_m(\mathbf{s}_k, t_i) + \widetilde{H}_m^k(t_i) + \epsilon_m^{k,i},$  
$$y_x^{k,i} = T_m(\mathbf{s}_k, t_i) + \widetilde{H}_{r,m}^k(t_i) + \frac{\widetilde{H}_{r,r}^k(t_i)}{2} T_r(\mathbf{s}_k, t_i) + \epsilon_x^{k,i},$$
 
$$y_n^{k,i} = T_m(\mathbf{s}_k, t_i) + \widetilde{H}_{r,m}^k(t_i) - \frac{\widetilde{H}_{r,r}^k(t_i)}{2} T_r(\mathbf{s}_k, t_i) + \epsilon_n^{k,i},$$





# **Modelling non-Gaussian quantities**

#### Power tail quantile (POQ) model

The quantile function  $F_{\theta}^{-1}(p)$ ,  $p \in [0, 1]$ , is defined through a quantile blend of left- and right-tailed generalised Pareto distributions:

$$f_{\theta}^{-}(p) = \begin{cases} \frac{1 - (2p)^{-\theta}}{2\theta}, & \theta \neq 0, \\ \frac{1}{2}\log(2p), & \theta = 0, \end{cases}$$

$$f_{\theta}^{+}(p) = -f_{\theta}^{-}(1-p) = \begin{cases} \frac{(2(1-p))^{-\theta}-1}{2\theta}, & \theta \neq 0, \\ -\frac{1}{2}\log(2(1-p)), & \theta = 0. \end{cases}$$

$$F_{\theta}^{-1}(p) = \theta_{0} + \frac{\tau}{2} \left[ (1-\gamma)f_{\theta_{3}}^{-}(p) + (1+\gamma)f_{\theta_{4}}^{+}(p) \right].$$

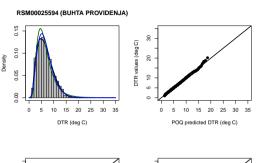
The parameters  $\theta = (\theta_0, \theta_1 = \log \tau, \theta_2 = \operatorname{logit}[(\gamma + 1)/2], \theta_3, \theta_4)$  control the median, spread/scale, skewness, and the left and right tail shape.

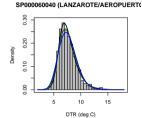
This model is also known as the *five parameter lambda model* (Gilchrist, 2000).

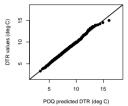


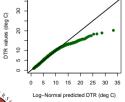


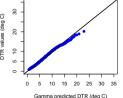
## **Diurnal range distributions**

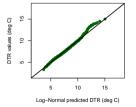


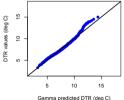










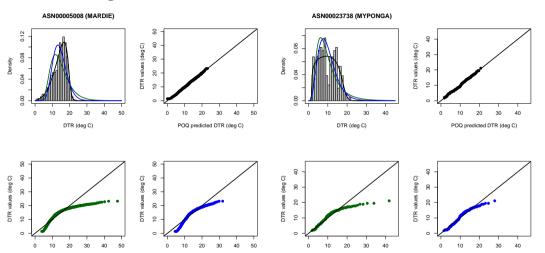




 $\lesssim$  For these stations, POQ does a slightly better job than a Gamma distribution.

## **Diurnal range distributions**

Log-Normal predicted DTR (deg C)



Log-Normal predicted DTR (deg C)

For these stations only POQ comes close to representing the distributions.

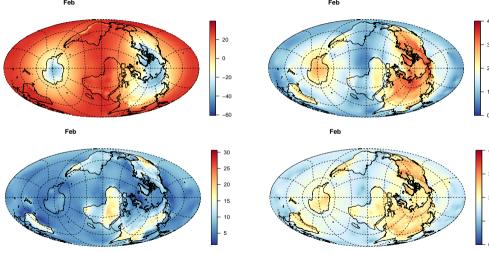
Note: Some shapes may be due to unmodeled station inhomogeneities.

Gamma predicted DTR (deg C)

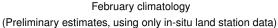


Gamma predicted DTR (deg C)

# Estimates of median & scale for $T_m$ and $T_r$









#### **Linearised inference**

All spatio-temporal latent random processes combined into  $x=(u,\beta,b)$ , with joint expectation  $\mu_x$  and precision  $Q_x$ :

$$egin{aligned} (m{x} \mid m{ heta}) &\sim \mathcal{N}(m{\mu}_x, m{Q}_x^{-1}) & ext{(Prior)} \ (m{y} \mid m{x}, m{ heta}) &\sim \mathcal{N}(h(m{x}), m{Q}_{y \mid x}^{-1}) & ext{(Observations)} \ p(m{x} \mid m{y}, m{ heta}) &\propto p(m{x} \mid m{ heta}) p(m{y} \mid m{x}, m{ heta}) & ext{(Conditional posterior)} \end{aligned}$$

#### Non-linear and/or non-Gaussian observations

For a non-linear h(x) with Jacobian J at  $x = \mu^*$ , iterate:

$$\begin{split} (\boldsymbol{x} \mid \boldsymbol{y}, \boldsymbol{\theta}) &\overset{\text{approx}}{\sim} \mathcal{N}(\widetilde{\boldsymbol{\mu}}, \widetilde{\boldsymbol{Q}}^{-1}) \qquad \text{(INLA posterior from } \overline{h}(\boldsymbol{x}) = h(\boldsymbol{\mu}^*) + \boldsymbol{J}(\boldsymbol{x} - \boldsymbol{\mu}^*) \text{)} \\ \widetilde{\boldsymbol{Q}} &= \boldsymbol{Q}_x + \boldsymbol{J}^\top \boldsymbol{Q}_{y|x} \boldsymbol{J} \qquad \text{(Generally: } \boldsymbol{Q}_x - \nabla_x \nabla_x^\top \log p(\boldsymbol{y}|\boldsymbol{x}, \boldsymbol{\theta}) \text{)} \\ \boldsymbol{\mu}^*_{\text{new}} &= \boldsymbol{\mu}^* + (\widetilde{\boldsymbol{\mu}} - \boldsymbol{\mu}^*) \cdot \operatorname*{argmin}_{a>0} \|\overline{h}(\widetilde{\boldsymbol{\mu}}) - h(\boldsymbol{\mu}^* + (\widetilde{\boldsymbol{\mu}} - \boldsymbol{\mu}^*)a) \| \end{split}$$





#### References

- Rue, H. and Held, L.: Gaussian Markov Random Fields; Theory and Applications; Chapman & Hall/CRC, 2005
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- Lindgren, F., Bolin, D., and Rue, H.: The SPDE Approach for Gaussian and Non-Gaussian Fields: 10 Years and Still Running; *Spatial Statistics, Special Issue: The Impact of Spatial Statistics*, 50:100599, 2022. https://arxiv.org/abs/2111.01084
- Lindgren, F., Haakon Bakka, David Bolin, Elias Krainski, Håvard Rue: A diffusion-based spatio-temporal extension of Gaussian Matérn fields, arXiv 2020–2023. https://arxiv.org/abs/2006.04917
- ▶ Video illustrating the results, produced by Philip Brohan: https://twitter.com/philipbrohan/status/1253411283598073867 https://player.vimeo.com/video/403663259
- Links to EUSTACE project reports and data:

  https://www.eustaceproject.org/

# Standardised observation uncertainty models

- Each data source may have complicated dependence structure
- To facilitate information blending, use a common error term structure

#### Common satellite derived data error model framework

The observational&calibration errors are modelled as three error components:

- $\blacktriangleright$  independent ( $\epsilon_0$ ),
- $\triangleright$  spatially and/or temporally correlated ( $\epsilon_1$ ), and
- systematic ( $\epsilon_2$ ),

with distributions determined by the uncertainty information from satellite calibration models.

E.g., 
$$y_i = T_m(\mathbf{s}_i, t_i) + \epsilon_0(\mathbf{s}_i, t_i) + \epsilon_1(\mathbf{s}_i, t_i) + \epsilon_2(\mathbf{s}_i, t_i)$$

In practice, each data source might have several different components of each type; independent components can be merged, but not necessarily correlated or systematic components.



