Parallelisable string backgrounds

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- hep-th/0308141, JHEP 10 (2003) 012

- hep-th/0305079, Class. Quant. Grav. 20 (2003) 3327-3340
- hep-th/0308141, JHEP **10** (2003) 012, with KAWANO Teruhiko and YAMAGUCHI Satoshi

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In this talk: type II string backgrounds from supergravity.

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$$\int_{M} e^{-2\phi} \left(R + 4|d\phi|^{2} - \frac{1}{2}|H|^{2} \right) d\text{vol}_{g}$$

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If M is simply-connected, this implies that M is parallelisable.

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ullet if M is simply-connected then flatness of D is also sufficient

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These geometries are easily characterised.

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[Chamseddine-FO-Sabra hep-th/0306278]

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[Cartan-Schouten (1926), Wolf (1970)]

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[Cartan-Schouten (1926), Wolf (1970)]

Only the first two have dH = 0.

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Summary

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Summary: Parallelisable geometries with closed torsion 3-form

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Summary: Parallelisable geometries with closed torsion 3-form are locally isometric to Lie groups with bi-invariant metrics.

Lie groups with bi-invariant metrics

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Which Lie algebras have an invariant metric?

abelian Lie algebras

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$$\mathfrak{h} o \mathfrak{so}(\mathfrak{g}) \cong \Lambda^2 \mathfrak{g}$$

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- we build the corresponding central extension $\mathfrak{g} \times_{\omega} \mathfrak{h}^*$
- \mathfrak{h} acts on $\mathfrak{g} \times_{\omega} \mathfrak{h}^*$ preserving the Lie bracket, so we can form the double extension

$$\mathfrak{d}(\mathfrak{g},\mathfrak{h})=\mathfrak{h}\ltimes(\mathfrak{g} imes_{\omega}\mathfrak{h}^*)$$

$$\mathfrak{g} \qquad \mathfrak{h} \qquad \mathfrak{h}^* \\
\mathfrak{g} \qquad \left(\langle -, - \rangle_{\mathfrak{g}} \quad 0 \quad 0 \\
0 \quad B \quad \text{id} \\
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[See also FO-Stanciu hep-th/9506152]

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(Any semisimple factors in \mathfrak{a} factor out of the double extension.

[FO-Stanciu hep-th/9402035])

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• we will call them CW(J)

All ten-dimensional lorentzian parallelisable spacetimes can be built out of:

Space Torsion

Space	Torsion
AdS_3	

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AdS_3	$dH = 0 H ^2 < 0$

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Ten-dimensional parallelisable geometries

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$$\begin{array}{lll} \operatorname{AdS}_{3} \times S^{7} & \operatorname{AdS}_{3} \times S^{3} \times \mathbb{R}^{4} \\ \operatorname{AdS}_{3} \times S^{3} \times S^{3} \times S^{3} & \operatorname{AdS}_{3} \times S^{3} \times \mathbb{R}^{7} \\ \operatorname{\mathbb{R}}^{1,0} \times S^{3} \times S^{3} \times S^{3} & \operatorname{\mathbb{R}}^{1,1} \times \operatorname{SU}(3) \\ \operatorname{\mathbb{R}}^{1,2} \times S^{7} & \operatorname{\mathbb{R}}^{1,3} \times S^{3} \times S^{3} \\ \operatorname{\mathbb{R}}^{1,6} \times S^{3} & \operatorname{\mathbb{R}}^{1,9} & \operatorname{CW}_{10}(J) \\ \operatorname{CW}_{6}(J) \times S^{3} \times \mathbb{R} & \operatorname{CW}_{8}(J) \times \mathbb{R}^{2} \\ \operatorname{CW}_{4}(J) \times S^{3} \times S^{3} & \operatorname{CW}_{4}(J) \times S^{3} \times S^{3} \\ \operatorname{CW}_{4}(J) \times \mathbb{R}^{6} & \operatorname{CW}_{4}(J) \times S^{3} \times \mathbb{R}^{3} \end{array}$$

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The case of linear dilaton was analysed by Kawano and Yamaguchi.

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For non-dilatonic backgrounds, this equation has solutions if and only if $|H|^2 = 0$; which restricts the possible geometries.

Spacetime	Supersymmetry
$AdS_3 \times S^3 \times S^3 \times \mathbb{R}$	16
$AdS_3 \times S^3 \times \mathbb{R}^4$	16
$CW_{10}(J)$	16,18(A),20,22(A),24(B),28(B)
$\mathrm{CW}_8(J) imes \mathbb{R}^2$	16,20
$\mathrm{CW}_6(J) \times \mathbb{R}^4$	16,24
$\mathrm{CW}_4(J) imes \mathbb{R}^6$	16
$\mathbb{R}^{1,9}$	32

$$AdS_3 \times S^3 \times S^3 \times \mathbb{R}$$

$$AdS_3 \times S^3 \times \mathbb{R}^4$$

$$\mathbb{R}^{1,1} \times SU(3)$$

$$\mathbb{R}^{1,3} \times S^3 \times S^3$$

$$\mathbb{R}^{1,6} \times S^3$$

$$\mathbb{R}^{1,9}$$

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$\mathbb{R}^{1,1} \times \mathrm{SU}(3)$	$\mathrm{CW}_6(J) imes \mathbb{R}^4$
$\mathbb{R}^{1,3} imes S^3 imes S^3$	$\mathrm{CW}_6(J) \times S^3 \times \mathbb{R}$
$\mathbb{R}^{1,6} \times S^3$	$\mathrm{CW}_4(J) \times S^3 \times \mathbb{R}^3$
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[Kawano-Yamaguchi hep-th/0306038]

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All these backgrounds are exact string backgrounds

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Non-simply connected backgrounds are obtained by orbifolding.

Type I supergravity coupled to supersymmetric Yang-Mills

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Parallelisable heterotic backgrounds

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- $dH = \frac{N}{2} \operatorname{Tr} F \wedge F + \cdots$
- action

$$\int_{M} e^{-2\phi} \left(R + 4|d\phi|^2 - \frac{1}{2}|H|^2 - \frac{N}{2}|F|^2 \right) d\text{vol}_g$$

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- $\bullet \ \delta^{D,A}(e^{-2\phi}F) = 0$

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Parallelisable backgrounds with ${\cal F}=0$

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Spacetime	Supersymmetry
$AdS_3 \times S^3 \times S^3 \times \mathbb{R}$	8
$AdS_3 \times S^3 \times \mathbb{R}^4$	8
$CW_{10}(J)$	8,10,12,14
$\mathrm{CW}_8(J) imes \mathbb{R}^2$	8,10
$\mathrm{CW}_6(J) imes \mathbb{R}^4$	8,12
$\mathrm{CW}_4(J) imes \mathbb{R}^6$	8
$\mathbb{R}^{1,9}$	16

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- $AdS_3 \times S^3 \times \mathbb{R}^4 \rightsquigarrow CW_6(J) \times \mathbb{R}^4$, and $AdS_3 \times S^3 \times S^3 \times \mathbb{R} \rightsquigarrow CW_8(J) \times \mathbb{R}^2$ by taking a Penrose limit

Thank you.